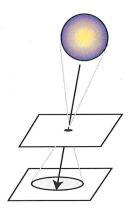
Student Guide to Activity 1: Features of the Sun



Problem: What do the following features look like on photographs of the Sun: sunspots, plages, solar flares, prominences, filaments, the corona, helmet streamers, and coronal holes? How do these features compare and contrast?



Caution: Never look directly at the Sun. To view the Sun, project an image through a card or sheet of notebook paper, pierced with pin sized hole, onto a sheet of white paper. The Sun's inverted image will appear on the paper below.

Introduction



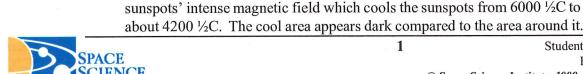
Our Sun is a middle-aged, medium sized star, big enough to hold a million Earths. The ancient Greeks thought that the Sun was a perfect sphere of fire. Today we know that the Sun is a variable (changeable) star that produces life giving light and heat as well as harmful radiation. It causes space weather that can harm astronauts working in space and can interfere with satellites orbiting our planet.

Features of the Sun's Surface and Atmosphere:

Although the average distance from

Earth to the Sun is a whopping 149,600,000 kilometers (93,000,000 miles), careful observation from Earth reveals a surprisingly large number of different visible features. The most obvious and best known feature is the sunspot. Typically moving in groups, these dark (in visible light), planet-sized features have been known to humankind for centuries. As sunspots form and disappear over periods of days or weeks, they also appear to move across the Sun's surface. Composed of strong magnetic fields, sunspots are shaped much like a horseshoe magnet that rises from below the Sun's surface. More accurately, however, flexible magnetic tubes, or "flux tubes," probably give rise to the

magnetic fields that we see. The rising hot gas is trapped by the







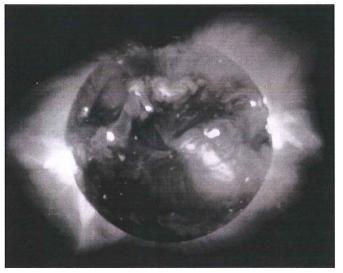
Thus, from Earth, we see spots on the Sun. In some photographs, we can also see light colored areas around groups of sunspots that resemble tufts of cotton candy. We call these fluffy looking fringes plages.



Sunspots are the source of massive releases of energy called **solar flares**, the most violent events in the solar system. In a matter of minutes to several hours, a solar flare releases about 10,000 times the annual energy consumption of the U.S. Solar flares give off radiation that includes X-rays, ultraviolet rays, and charged particles called protons and electrons. This sudden surge in radiation can damage spacecraft and even give a dose or radiation to travelers flying in airplanes over the polar regions.



Also visible for only minutes, are granulations in the Sun's photosphere. Granulations are rising and falling columns of hot gases that look like fluffy marshmallows arranged in a honeycomb pattern. The tops of these granules form the Sun's "surface." Although we refer to the Sun's "surface" as the photosphere, you probably know that the Sun has no solid surface, unlike Earth. It is an uneven sphere of glowing, hot gas!



Courtesy of the Lockheed Palo Alto Research Laboratory and the National Astronomical Observatory in Japan.

The bright areas in the X-ray image from the Japanese Yohkoh satellite are called "active regions," which contain hot, dense gas. They are also the source of the most intense X-ray eminations. The dark areas are coronal holes.



Just as the Sun disappears behind the Moon during a total solar eclipse, a flash of bright red light appears. This colorful layer of the Sun, called the chromosphere, becomes visible for a brief instant. Although we know little about the chromosphere, there are curious, permanent features of the chromosphere, called spicules, that we can study in more detail. There are so many of these fine, bright, hairlike features, that they are always visible near the Sun's edge, even though an individual spicule lasts only minutes. Like sunspots, spicules rise and fall vertically above the Sun's surface.



One of the most spectacular features of the Sun are solar **prominences**. They appear to stream, loop and arch away from the Sun. The most recognizable prominences appear as huge arching columns of gas above the limb (edge) of the Sun. However, when prominences are photographed on the surface of the Sun, they appear as long, dark, threadlike objects and are called **filaments**. Like sunspots, prominences are cooler (about 10,000 ½C) in relation to the much hotter background of the Sun's outer atmosphere (about 1,500,000 ½C). Prominences can also erupt from the Sun with a tremendous burst of energy.







If you have seen photographs of a solar eclipse, then you have probably noticed a bright halo around the Sun, called the **corona**. Sometimes parts of the corona appear to be missing. Logically, we call this area a **coronal hole**. Scientists believe that the **solar** wind, a million mile per hour gale that blows away from the Sun, originates in coronal holes. Unlike wind on Earth, the solar wind is a stream of ionized (electrically charged) particles speeding away from the Sun.



The Sun's corona changes with sunspot activity. When there are more sunspots, the corona appears to be held closely to the Sun; when there are fewer sunspots, the corona streams out into space in a shape that resembles the spike on a warlike, peaked helmet called **helmet streamers**. While helmet streamers are long-lived, their demise often occurs abruptly through massive and powerful eruptions called coronal mass ejections (CMEs).



Artist's conception of a coronal mass ejection moving away from the Sun toward Earth.



These huge clouds of hot solar gas and magnetic fields are often associated with solar flares. They can cause **magnetic storms** when they hit Earth's magnetic field and damage human technological systems in space and on the ground. For example, in 1989, the Quebec province in Canada suffered an electrical blackout because many transformers were destroyed by a large magnetic storm. That one storm caused many millions of dollars worth of damage. A powerful solar flare erupted from the Sun about three days before the start of the storm at Earth. Even when the Sun is not too active, solar storms can cause problems. A magnetic storm on January 11, 1997 was blamed for the loss of a \$270 million dollar AT&T communications satellite. This moderate storm was caused by a coronal mass ejection that erupted from the Sun even though there were no noticeable sunspots.





ACTIVITY 1 STUDENT WORKSHEET- Features of	f the Sun Name:
Goal:	
Procedure:	
Color, label, and list any features that	ky note paper to draw a diagram of the sun. t you know of. When you have decided on a the teacher to come over and approve.
 After all members have finished with directions for the guided highlight for of your table. 	the drawing, group member #1 will read the the group. Directions are in the bin in the middle
this time including the following featu	p is going to create a drawing of the sun again, ures (on another large sticky note) First draw a res using your highlighted information, color, and
SunspotPlagesSolar flare	SpiculesProminences
Granulations	 Filaments
PhotosphereChromosphere	CoronaCoronal hole

4.	How does your new picture compare to your group's first picture?	
5.	After you have completed the wisdom walk, make any changes that you would like t your picture. Once you have agreed upon a model, draw your labeled diagram below.	0

For each of the following features of the sun, write a short description the feature's appearance and a short definition.

Plages Solar flare Granulations Photosphere Chromosphere Spicules Frominences Frominences Coronal Coronal hole	Feature	Definition	Description